

DISCRETE QUANTUM GRAVITY and QUANTUM FIELD THEORY

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Abstract

We introduce a discrete 4-dimensional module over the integers that appears to have maximal symmetry. By adjoining the usual Minkowski distance, we obtain a discrete 4-dimensional Minkowski space. Forming universe histories in this space and employing the standard causal order, the histories become causal sets. These causal sets increase in size rapidly and describe an inflationary period for the early universe. We next consider the symmetry group G for the module. We show that G has order 24 and we construct its group table. In a sense G is a discrete approximation to the Lorentz group. However, we note that it contains no boosts and is essentially a rotation group. Unitary representations of G are constructed. The energy-momentum space dual to the discrete module is obtained and a quantum formalism is derived. A discrete Fock space is introduced on this structure and free quantum fields are considered. Finally, we take the first step in a study of interacting quantum fields.

1 Introduction

It is well-known that general relativity and quantum field theory are both plagued by singularities and infinities. This is particularly serious for quantum field theory because small distances must be considered there. These difficulties are usually circumvented by methods of infinity cancellations and re-normalizations but the methods are mathematically suspect. In general relativity, it is usually just admitted that the theory breaks down and is no longer applicable at small distance scales. The simplest and possibly only solution to these problems has been known for almost a hundred years [1, 5, 6]. At that time, it was suggested by several researchers that the physical universe is discrete. They speculated that there exists an elementary length ℓ and an elementary time t . The likely values of these are the Planck length $\ell \approx 10^{-33}\text{cm}$ and the Planck time $t \approx 10^{-43}\text{sec}$. In this view, spacetime is composed of discrete, tiny cells of Planck size. This idea was not completely outrageous because it was already known that energies were composed of packets which came in multiples of Planck's constant h and electric charge came in multiples of the electron charge e . (The latter was later altered when quarks were discovered with charges $\pm 1/3$ and $\pm 2/3$, but the idea is still the same.) One problem is that, unlike energy and charge, this granular structure of spacetime has not been experimentally observed. However, some investigators believe that with ever more sensitive instruments, this discrete framework will eventually be unveiled.

The main problem with discrete spacetime is that a substantial amount of symmetry would be lost. In particular, we would lose Lorentz invariance. Lorentz invariance is not only a pillar of theoretical physics, it has been experimentally verified a myriad of times. But it is possible that Lorentz invariance is only an approximation. There may be a smaller, more fundamental symmetry group that is indistinguishable from the Lorentz group except at very small scales.

This article attempts to construct a reasonable discrete spacetime upon which a discrete quantum gravity and quantum field theory can be built. How should one proceed with such a construction? Let's start with 2-dimensional space and add time later. We could naively begin with a square lattice structure. Basic cells that we call *vertices* would be placed at $\hat{0} = (0, 0)$, $\hat{e}_1 = (1, 0)$, $\hat{e}_2 = (0, 1)$ and other vertices would have locations $a\hat{e}_1 + b\hat{e}_2$, $a, b \in \mathbb{Z}$. In this case, we set Planck's length $\ell = 1$. Now $\hat{0}$ is indeed a distance 1 from its four nearest neighbors $\pm\hat{e}_1, \pm\hat{e}_2$. But the distance from $\hat{0}$

to neighbor $\widehat{e}_1 + \widehat{e}_2$ is $\sqrt{2}$ and this is not a measurable distance which would have to be a positive integer. Also, \widehat{e}_1 and \widehat{e}_2 are distance $\sqrt{2}$ apart. The symmetry group for this square lattice is the group of order 4 consisting of the identity and the three rotations by radian angles $\pi/2$, π and $3\pi/2$.

We obtain more symmetry by considering a triangular lattice. In this case we have the basic vertices at $\widehat{0} = (0, 0)$, $\widehat{e} = (1, 0)$, $\widehat{f} = (1/2, \sqrt{3}/2)$ and other vertices have locations at $a\widehat{e} + b\widehat{f}$, $a, b \in \mathbb{Z}$. Now $\widehat{0}$ is a distance 1 from its six nearest neighbors $\pm\widehat{e}$, $\pm\widehat{f}$, $\widehat{e} - \widehat{f}$, $\widehat{f} - \widehat{e}$. We also have the bonus that \widehat{e} and \widehat{f} as well as \widehat{e} and $\widehat{e} - \widehat{f}$ and others are distance 1 apart. Unfortunately, the distances from \widehat{e} to $\widehat{f} - \widehat{e}$ is $\sqrt{3}$ which is not elementary length measurable. The triangular lattice has more symmetry than the square lattice because its symmetry group has order 6 and consists of the rotations by angles $n\pi/3$, $n = 0, 1, \dots, 5$.

We noted that the distances between many lattice vertices in both cases are not elementary length measurable because they are not integers. In the first case, the distance has the form

$$\|a\widehat{e}_1 + b\widehat{e}_2\|_2 = \sqrt{a^2 + b^2}$$

while in the second case

$$\|a\widehat{e} + b\widehat{f}\|_2 = \sqrt{a^2 + b^2 + ab}$$

However, this is not really a problem because as usual in relativity theory, we measure distance with the Minkowski metric which is in terms of the Euclidean distance squared so instead we have

$$\|a\widehat{e}_1 + b\widehat{e}_2\|_2^2 = a^2 + b^2$$

and

$$\|a\widehat{e} + b\widehat{f}\|_2^2 = a^2 + b^2 + ab$$

which, of course, are integers. Notice that these two cases are the only ones with this property. Indeed, suppose that \widehat{g} and \widehat{h} are unit vectors in \mathbb{R}^2 with inner product $\langle \widehat{g}, \widehat{h} \rangle = \cos \theta$, where $0 < \theta \leq \pi/2$. Then for $a, b \in \mathbb{Z}$ we have

$$\|a\widehat{g} + b\widehat{h}\|_2^2 = a^2 \|\widehat{g}\|_2^2 + b^2 \|\widehat{h}\|_2^2 + 2ab \langle \widehat{g}, \widehat{h} \rangle$$

$$= a^2 + b^2 + 2ab \cos \theta$$

Notice that the only values of θ for which this is always an integer are $\theta = \pi/3$ or $\pi/2$.

A similar analysis holds for 3-dimensional space. In this case we have a cubic or tetrahedral lattice. As in the 2-dimensional case, we shall see that the latter has certain advantages. Moreover, these are the only two that have measurable Minkowski distances.

In Section 2 we discuss discrete 3-dimensional space. We first construct the tetrahedral lattice \mathcal{S}_3 and its corresponding symmetry group $\mathcal{G}(\mathcal{S}_3)$. We show that $\mathcal{G}(\mathcal{S}_3)$ has order 24 and we exhibit its group table, In Section 3 we derive 2- and 3-dimensional unitary representations for $\mathcal{G}(\mathcal{S}_3)$.

Section 4 introduces a discrete 4-dimensional module $\mathcal{S}_4 \subseteq \mathbb{R}^4$ that has maximal symmetry in the sense we previously discussed. Letting $\|u\|_4$ be the usual Minkowski metric, the pair $(\mathcal{S}_4, \|\cdot\|_4)$ becomes a discrete 4-dimensional Minkowski space. In a sense, its symmetry group $\mathcal{G}(\mathcal{S}_4)$ is a discrete approximation to the Lorentz group. However, we note that $\mathcal{G}(\mathcal{S}_4)$ contains no boosts and is essentially a discrete rotation group. We next form universe histories \mathcal{S}^{t-} , $t = 1, 2, \dots$, where t represents a discrete time. Employing the standard causal order $u < v$ if $u^0 < v^0$ and $\|u - v\|_4^2 \geq 0$, \mathcal{S}^{t-} becomes a causal set (causet) [2–4].

The causets \mathcal{S}^{t-} increase rapidly in size and describe an inflationary period for the early universe. During this period, the universe is essentially flat and gravity does not present itself. Moreover, the causets have a special property that we call weak covariance. This property states that all paths between two fixed vertices have the same length. At the end of the inflationary period, the system experiences a phase transition and enters the multiverse period. During the multiverse period, the universe splits into parts, each with their own geometry. The various geometries then determine curvatures and gravity in the constituent universes. Moreover, the causets possess a stronger property called covariance [2–4].

The structure presented in Section 4 entails a novel phenomenon that is more fundamental than the constancy of the speed of light c in a vacuum. The reason that c is the upper speed limit is that c is the *only* speed that a particle can attain other than zero. The reason that we observe slower speeds is that we are measuring average speeds and c is the instantaneous speed of a particle.

In Section 5 the corresponding energy-momentum space $\widehat{\mathcal{S}}_4$ is obtained

and a quantum formalism on $\widehat{\mathcal{S}}_4$ is derived. A discrete Fock space is introduced on $\widehat{\mathcal{S}}_4$ and free quantum fields are studied. A first step toward interacting quantum fields is presented in Section 6 and it is mentioned that this framework may result in a mathematically rigorous quantum field theory.

2 Discrete Space

For contrast and comparison, we begin with a discrete space formed from a cubic lattice. To discuss this lattice, let $\widehat{e}_1 = (1, 0, 0)$, $\widehat{e}_2 = (0, 1, 0)$, $\widehat{e}_3 = (0, 0, 1)$ be the usual basic vertices and let $\mathcal{S}_c^3 \subseteq \mathbb{R}^3$ be the set of vertices

$$\mathcal{S}_c^3 = \{u \in \mathbb{R}^3 : u = n\widehat{e}_1 + p\widehat{e}_2 + q\widehat{e}_3, \quad n, p, q \in \mathbb{Z}\}$$

Notice that \mathcal{S}_c^3 is a module over the integers \mathbb{Z} . That is, $u, v \in \mathcal{S}_c^3$ implies $u + v \in \mathcal{S}_c^3$ and $nu \in \mathcal{S}_c^3$ for all $n \in \mathbb{Z}$. Now any $u \in \mathcal{S}_c^3$ is distance 1 to its six nearest neighbors $u \pm \widehat{e}_1$, $u \pm \widehat{e}_2$, $u \pm \widehat{e}_3$. We think of the vertices in \mathcal{S}_c^3 as cells of Planck size that may or may not be occupied by a particle and the six edges from u to its nearest neighbors as directions along which particles can move.

A *symmetry* on \mathcal{S}_c^3 is a linear isometry on \mathcal{S}_c^3 with determinant 1. We denote the group of symmetries on \mathcal{S}_c^3 by $\mathcal{G}(\mathcal{S}_c^3)$. The matrices of elements in $\mathcal{G}(\mathcal{S}_c^3)$ are determined by how they act on the basis $\widehat{e}_1, \widehat{e}_2, \widehat{e}_3$. For example, the symmetry

$$\widehat{A}: (\widehat{e}_1, \widehat{e}_2, \widehat{e}_3) \rightarrow (\widehat{e}_2, -\widehat{e}_1, \widehat{e}_3)$$

is represented by the matrix

$$\widehat{A} = \begin{bmatrix} 0 & -1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

It turns out that $\mathcal{G}(\mathcal{S}_c^3)$ is a group of order 24. These matrices are unitary with determinant 1. Besides \widehat{A} given above, the elements of $\mathcal{G}(\mathcal{S}_c^3)$ have matrices given as follows:

$$\widehat{B} = \begin{bmatrix} -1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 1 \end{bmatrix}, \quad \widehat{C} = \begin{bmatrix} 0 & 1 & 0 \\ -1 & 0 & 0 \\ 0 & 0 & 1 \end{bmatrix}, \quad \widehat{D} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & -1 \\ 0 & 1 & 0 \end{bmatrix}, \quad \widehat{E} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & -1 \end{bmatrix}$$

$$\begin{aligned}
\widehat{F} &= \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & -1 & 0 \end{bmatrix}, \widehat{G} = \begin{bmatrix} 0 & 0 & 1 \\ 0 & 1 & 0 \\ -1 & 0 & 0 \end{bmatrix}, \widehat{H} = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{bmatrix}, \widehat{I} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix} \\
\widehat{J} &= \begin{bmatrix} 0 & 0 & -1 \\ 0 & 1 & 0 \\ 1 & 0 & 0 \end{bmatrix}, \widehat{K} = \begin{bmatrix} 0 & 0 & 1 \\ 1 & 0 & 0 \\ 0 & 1 & 0 \end{bmatrix}, \widehat{L} = \begin{bmatrix} 0 & 1 & 0 \\ 0 & 0 & 1 \\ 1 & 0 & 0 \end{bmatrix}, \widehat{M} = \begin{bmatrix} 0 & -1 & 0 \\ 0 & 0 & -1 \\ 1 & 0 & 0 \end{bmatrix} \\
\widehat{N} &= \begin{bmatrix} 0 & 0 & 1 \\ -1 & 0 & 0 \\ 0 & -1 & 0 \end{bmatrix}, \widehat{O} = \begin{bmatrix} 0 & -1 & 0 \\ 0 & 0 & 1 \\ -1 & 0 & 0 \end{bmatrix}, \widehat{P} = \begin{bmatrix} 0 & 0 & -1 \\ -1 & 0 & 0 \\ 0 & 1 & 0 \end{bmatrix}, \widehat{Q} = \begin{bmatrix} 0 & 1 & 0 \\ 0 & 0 & 1 \\ 1 & 0 & 0 \end{bmatrix} \\
\widehat{R} &= \begin{bmatrix} 0 & 1 & 0 \\ 0 & 0 & -1 \\ -1 & 0 & 0 \end{bmatrix}, \widehat{S} = \begin{bmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & -1 \end{bmatrix}, \widehat{T} = \begin{bmatrix} 0 & 0 & 1 \\ 0 & -1 & 0 \\ 1 & 0 & 0 \end{bmatrix}, \widehat{U} = \begin{bmatrix} 0 & -1 & 0 \\ -1 & 0 & 0 \\ 0 & 0 & -1 \end{bmatrix} \\
\widehat{V} &= \begin{bmatrix} 0 & 0 & -1 \\ 0 & -1 & 0 \\ -1 & 0 & 0 \end{bmatrix}, \widehat{W} = \begin{bmatrix} 0 & 0 & -1 \\ 1 & 0 & 0 \\ 0 & -1 & 0 \end{bmatrix}, \widehat{X} = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{bmatrix}
\end{aligned}$$

We shall not write down the group table for $\mathcal{G}(\mathcal{S}_c^3)$ because we want to proceed to the tetrahedral space lattice which we believe has certain advantages. Let e, f, g be unit vectors in \mathbb{R}^3 that are distance 1 from each other. That is,

$$\|e\| = \|f\| = \|g\| = \|e - f\| = \|e - g\| = \|f - g\| = 1$$

An example of such vectors is

$$e = (1, 0, 0), \quad f = (1/2, \sqrt{3}/2, 0), \quad g = (1/2, 1/2\sqrt{3}, \sqrt{2/3}) \quad (2.1)$$

but the particular form is not needed now. The four vectors $(0, e, f, g)$ are distance 1 from each other and form the vertices of a tetrahedron with edge length 1. Notice that

$$\langle e, f \rangle = \langle e, g \rangle = \langle f, g \rangle = 1/2$$

so the angle between any two of e, f, g is $\pi/3$. As mentioned in Section 1, the cubic and tetrahedral space lattices are the only regular lattices for which the distance squared between any two vertices is an integer. The vertices of this lattice is given by the set

$$\mathcal{S}_3 = \{u \in \mathbb{R}^3 : u = ne + pf + qg, \quad n, p, q \in \mathbb{Z}\}$$

As before \mathcal{S}_3 is a module over \mathbb{Z} . One advantage of \mathcal{S}_3 over \mathcal{S}_c^3 is that each $u \in \mathcal{S}_3$ has 12 nearest neighbors a distance 1 away so we now have 12 directions along which a particle can propagate.

Lemma 2.1. (i) *The vectors e, f, g are linearly independent and form a basis for \mathcal{S}_3 . (ii) There are 12 unit vectors in \mathcal{S}_3 and these are given by $e, f, g, e - f, e - g, f - g$ and their negatives.*

Proof. (i) If $ne + pf + qg = 0$, taking inner products with e, f and g give

$$n + \frac{1}{2}p + \frac{1}{2}q = \frac{1}{2}n + p + \frac{1}{2}q = \frac{1}{2}n + \frac{1}{2}p + q = 0$$

Solving these equations simultaneously we have $n = p = q = 0$ so e, f, g are linear independent. Since $\dim \mathbb{R}^3 = 3$, $\{e, f, g\}$ forms a basis for \mathbb{R}^3 and hence a basis for \mathcal{S}_3 . (ii) If $u = ne + pf + qg$, $n, p, q \in \mathbb{Z}$, is a unit vector, we have that

$$1 = \|u\|^2 = n^2 + p^2 + q^2 + np + nq + pq = \frac{1}{2} [(n + p)^2 + (n + q)^2 + (p + q)^2]$$

Hence,

$$(n + p)^2 + (n + q)^2 + (p + q)^2 = 2 \tag{2.2}$$

The only way that (2.2) can hold is if two of the terms on the left side of (2.2) are 1 and the other is 0. This is possible if and only if two of the numbers n, p, q are 0 and the third is ± 1 or if one of the numbers is 1, another is -1 and the third is 0. This gives the 12 possibilities listed above. \square

A *triad* is three unit vectors in \mathcal{S}_3 whose inner products are $1/2$. There are 8 triads given by: $\{e, f, g\}, \{-e, -f, -g\}, \{e, e - f, e - g\}, \{-e, f - e, g - e\}, \{f, f - e, f - g\}, \{-f, e - f, g - f\}, \{g, g - e, g - f\}, \{-g, e - g, f - g\}$. Each triad corresponds to three *triples* written in cyclic order. For example, corresponding to triad $\{e, f, g\}$ we have the triples $(e, f, g), (f, g, e), (g, e, f)$. We call (e, f, g) the *basic triple*. A *symmetry* on \mathcal{S}_3 is a linear transformation $T: \mathcal{S}_3 \rightarrow \mathcal{S}_3$ that takes triples to triples and has determinant 1. The

symmetries are determined by their action on the basic triple. For example if $Te = f$, $Tf = f - e$ and $Tg = f - g$, then T preserves all triples. For instance,

$$\begin{aligned} T(-g) &= g - f \\ T(e - g) &= Te - Tg = f - f + g = g \\ T(f - g) &= Tf - Tg = f - e - f + g = g - e \end{aligned}$$

Hence, $T(-g, e - g, f - g) = (g - f, g, g - e)$.

Since there are 24 triples, we conclude that there are 24 symmetries. The simplest are the identity $I(e, f, g) = (e, f, g)$ and $A(e, f, g) = (f, g, e)$. We can write the symmetries relative to the basis $\{e, f, g\}$ as follows:

$$\begin{aligned} A &= \begin{bmatrix} 0 & 0 & 1 \\ 1 & 0 & 0 \\ 0 & 1 & 0 \end{bmatrix}, B = \begin{bmatrix} 0 & 1 & 0 \\ 0 & 0 & 1 \\ 1 & 0 & 0 \end{bmatrix}, C = \begin{bmatrix} -1 & -1 & -1 \\ 0 & 1 & 0 \\ 1 & 0 & 0 \end{bmatrix}, D = \begin{bmatrix} 0 & 0 & 1 \\ 0 & 1 & 0 \\ -1 & -1 & -1 \end{bmatrix} \\ E &= \begin{bmatrix} 1 & 0 & 0 \\ -1 & -1 & -1 \\ 0 & 1 & 0 \end{bmatrix}, F = \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ -1 & -1 & -1 \end{bmatrix}, G = \begin{bmatrix} -1 & -1 & -1 \\ 1 & 0 & 0 \\ 0 & 0 & 1 \end{bmatrix}, H = \begin{bmatrix} 0 & 1 & 0 \\ -1 & -1 & -1 \\ 0 & 0 & 1 \end{bmatrix} \\ I &= \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}, J = \begin{bmatrix} -1 & -1 & -1 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{bmatrix}, K = \begin{bmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ -1 & -1 & -1 \end{bmatrix}, L = \begin{bmatrix} 0 & 0 & 1 \\ -1 & -1 & -1 \\ 1 & 0 & 0 \end{bmatrix} \\ M &= \begin{bmatrix} 1 & -1 & 1 \\ 0 & -1 & 0 \\ 0 & 0 & -1 \end{bmatrix}, N = \begin{bmatrix} 0 & 0 & -1 \\ 1 & 1 & 1 \\ 0 & -1 & 0 \end{bmatrix}, O = \begin{bmatrix} 0 & -1 & 0 \\ 0 & 0 & -1 \\ 1 & 1 & 1 \end{bmatrix}, P = \begin{bmatrix} -1 & 0 & 0 \\ 0 & -1 & 0 \\ 1 & 1 & 1 \end{bmatrix} \\ Q &= \begin{bmatrix} 0 & 0 & -1 \\ 0 & -1 & 0 \\ -1 & 0 & 0 \end{bmatrix}, R = \begin{bmatrix} 1 & 1 & 1 \\ -1 & 0 & 0 \\ 0 & -1 & 0 \end{bmatrix}, S = \begin{bmatrix} 1 & 1 & 1 \\ 0 & 0 & -1 \\ -1 & 0 & 0 \end{bmatrix}, T = \begin{bmatrix} -1 & 0 & 0 \\ 1 & 1 & 1 \\ 0 & 0 & -1 \end{bmatrix} \end{aligned}$$

$$U = \begin{bmatrix} 0 & -1 & 0 \\ -1 & 0 & 0 \\ 0 & 0 & -1 \end{bmatrix}, \quad V = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 0 & -1 \\ 0 & -1 & 0 \end{bmatrix}, \quad W = \begin{bmatrix} 0 & -1 & 0 \\ 1 & 1 & 1 \\ -1 & 0 & 0 \end{bmatrix}, \quad X = \begin{bmatrix} 0 & 0 & -1 \\ -1 & 0 & 0 \\ 1 & 1 & 1 \end{bmatrix}$$

We denote the group of symmetries $\{A, B, \dots, X\}$ by $\mathcal{G}(\mathcal{S}_3)$. There are other linear transformations on \mathcal{S}_3 that take triples to triples but these do not have unit determinant. For example,

$$A' = \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{bmatrix}$$

has $\det A = -1$. Even though $\mathcal{G}(\mathcal{S}_c^3)$ and $\mathcal{G}(\mathcal{S}_3)$ both have order 24, they are different groups and we believe that $\mathcal{G}(\mathcal{S}_3)$ is more interesting. The group table for $\mathcal{G}(\mathcal{S}_3)$ is given by Table 1.

Some of the subgroups of $\mathcal{G}(\mathcal{S}_3)$ are: $\{I, A, B\}$, $\{I, C, D\}$, $\{I, E, F\}$, $\{I, G, H\}$, $\{I, W, J, X\}$, $\{I, A, B, U, V, Q\}$, $\{A, B, \dots, L\}$, $\{I, J\}$, $\{I, K\}$, $\{I, L\}$, $\{I, T\}$, $\{I, U\}$, $\{I, V\}$, $\{I, M\}$. We can check from Table 1 that M and N generate $\mathcal{G}(\mathcal{S}_3)$. In fact, any two non-commuting elements from $\{M, N, \dots, X\}$ are generators. As usual, a linear transformation $Z: \mathcal{S}_3 \rightarrow \mathcal{S}_3$ is an *isometry* if $\langle Zu, Zv \rangle = \langle u, v \rangle$ for all $u, v \in \mathcal{S}_3$. It is clear that an isometry with unit determinant is a symmetry. Conversely, it is easy to check that M and N are isometries and since M and N generate $\mathcal{G}(\mathcal{S}_3)$, it follows that every element of $\mathcal{G}(\mathcal{S}_3)$ is an isometry. We conclude that our definition of a symmetry on \mathcal{S}_3 coincides with our original definition of a symmetry.

	I	A	B	C	D	E	F	G	H	J	K	L	M	N	O	P	Q	R	S	T	U	V	W	X
I	I	A	B	C	D	E	F	G	H	J	K	L	M	N	O	P	Q	R	S	T	U	V	W	X
A	A	B	I	E	J	K	G	L	D	H	C	F	N	O	M	R	V	W	T	X	Q	U	P	S
B	B	I	A	K	H	C	L	F	J	D	E	G	O	M	N	W	U	P	X	S	V	Q	R	T
C	C	G	J	D	I	L	B	K	E	F	A	H	P	T	V	Q	M	X	O	W	R	S	N	U
D	D	K	F	I	C	H	J	A	L	B	G	E	Q	W	S	M	P	U	V	N	X	O	T	R
E	E	L	H	J	A	F	I	C	K	G	B	D	R	X	U	V	N	S	M	P	W	T	O	Q
F	F	D	K	G	L	I	E	J	B	C	H	A	S	Q	W	T	X	M	R	V	O	P	U	N
G	G	J	C	L	F	A	K	H	I	E	D	B	T	V	P	X	S	N	W	U	M	R	Q	O
H	H	E	L	B	K	J	D	I	G	A	F	C	U	R	X	O	W	V	Q	M	T	N	S	P
J	J	C	G	A	E	D	H	B	F	I	L	K	V	P	T	N	R	Q	U	O	S	M	X	W
K	K	F	D	H	B	G	A	E	C	L	I	J	W	S	Q	U	O	T	N	R	P	X	M	V
L	L	H	E	F	G	B	C	D	A	K	J	I	X	U	R	S	T	O	P	Q	N	W	V	M
M	M	R	S	Q	P	N	O	U	T	V	X	W	I	E	F	D	C	A	B	H	G	J	L	K
N	N	W	T	V	R	O	M	Q	X	U	S	P	A	K	G	J	E	B	I	D	L	H	F	C
O	O	P	X	U	W	M	N	V	S	Q	T	R	B	C	L	H	K	I	A	J	F	D	G	E
P	P	X	O	M	Q	T	V	R	W	S	U	N	C	L	B	I	D	G	J	E	K	F	H	A
Q	Q	U	V	P	M	W	S	X	N	O	R	T	D	H	J	C	I	K	F	L	A	B	E	G
R	R	S	M	N	V	X	U	W	P	T	Q	O	E	F	I	A	J	L	H	K	C	G	D	B
S	S	M	R	X	T	Q	W	O	V	P	N	U	F	I	E	L	G	D	K	B	J	C	A	H
T	T	N	W	S	X	V	P	M	U	R	O	Q	G	A	K	F	L	J	C	I	H	E	B	D
U	U	V	Q	W	O	R	X	T	M	N	P	S	H	J	D	K	B	E	L	G	I	A	C	F
V	V	Q	U	R	N	P	T	S	O	M	W	X	J	D	H	E	A	C	G	F	B	I	K	L
W	W	T	N	O	U	S	Q	P	R	X	V	M	K	G	A	B	H	F	D	C	E	L	J	I
X	X	O	P	T	S	U	R	N	Q	W	M	V	L	B	C	G	F	H	E	A	D	K	I	J

Table 1

3 Unitary Representations

This section constructs unitary representations for the group $\mathcal{G}(\mathcal{S}_3)$. The matrix realizations that we gave for $\mathcal{G}(\mathcal{S}_3)$ were not necessarily unitary because they were given relative to the nonorthogonal basis $\{e, f, g\}$. Since these matrices are isometries, if we represent them in the standard basis $\hat{e}_1 = (1, 0, 0)$, $\hat{e}_2 = (0, 1, 1)$, $\hat{e}_3 = (0, 0, 1)$, they will become unitary. Using the concrete form (2.1) for e, f, g , the two bases are related by

$$e = \hat{e}_1, \quad f = \frac{1}{2}\hat{e}_1 + \frac{\sqrt{3}}{2}\hat{e}_2, \quad g = \frac{1}{2}\hat{e}_1 + \frac{1}{2\sqrt{3}}\hat{e}_2 + \sqrt{\frac{2}{3}}\hat{e}_3$$

and

$$\hat{e}_1 = e, \quad \hat{e}_2 = -\frac{1}{\sqrt{3}}e + \frac{2}{\sqrt{3}}f, \quad \hat{e}_3 = -\frac{1}{\sqrt{6}}e - \frac{1}{\sqrt{6}}f + \sqrt{\frac{3}{2}}g$$

The basis transformations become:

$$U = \begin{bmatrix} 1 & -1/\sqrt{3} & -1/\sqrt{6} \\ 0 & 2/\sqrt{3} & -1/\sqrt{6} \\ 0 & 0 & \sqrt{3}/2 \end{bmatrix}, \quad U^{-1} = \begin{bmatrix} 1 & 1/2 & 1/2 \\ 0 & \sqrt{3}/2 & 1/2\sqrt{3} \\ 0 & 0 & \sqrt{2/3} \end{bmatrix}$$

Define the unitary representation \mathcal{U} of $\mathcal{G}(\mathcal{S}_3)$ on \mathbb{C}^3 by

$$\mathcal{U}(Z) = U^{-1}ZU$$

for all $Z \in \mathcal{G}(\mathcal{S}_3)$. This is a group representation because

$$\mathcal{U}(YZ) = U^{-1}YZU = U^{-1}YUU^{-1}ZU = \mathcal{U}(Y)\mathcal{U}(Z)$$

Since M and N generate $\mathcal{G}(\mathcal{S}_3)$ we find $\mathcal{U}(M), \mathcal{U}(N)$ to be:

$$\mathcal{U}(M) = \begin{bmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & -1 \end{bmatrix}, \quad \mathcal{U}(N) = \begin{bmatrix} 1/2 & -1/2\sqrt{3} & -\sqrt{2/3} \\ \sqrt{3}/2 & 1/6 & \sqrt{2}/3 \\ 0 & -2\sqrt{2}/3 & 1/3 \end{bmatrix}$$

Clearly, $\mathcal{U}(M)$ is a unitary and it is easy to check that $\mathcal{U}(N)$ is unitary. In fact, they are orthogonal matrices because they have real entries. Since M

and N generate $\mathcal{G}(\mathcal{S}_3)$, it again follows that $\mathcal{U}(Z)$ is unitary for all $Z \in \mathcal{G}(\mathcal{S}_3)$ so \mathcal{U} is a unitary representation. Since $\mathcal{G}(\mathcal{S}_3)$ and $\mathcal{U}[\mathcal{G}(\mathcal{S}_3)]$ are isomorphic groups, we can and frequently will identify them.

All the matrices in $\mathcal{G}(\mathcal{S}_3)$ have eigenvalues among the numbers $\pm 1, \pm i, e^{\pm 2\pi i/3}$ with possible multiplicities. Of course, their corresponding eigenvectors are different, in general. For example, A has eigenvalues $1, e^{2\pi i/3}, e^{-2\pi i/3}$ with corresponding (unnormalized) eigenvectors $(1, 1, 1), (e^{2\pi i/3}, 1, e^{-2\pi i/3}), (e^{-2\pi i/3}, 1, e^{2\pi i/3})$ relative to the $\{e, f, g\}$ basis. It is easy to check that B, C, \dots, H have the same eigenvalues as A . A new pattern begins with J which has eigenvalues $1, -1$ (multiplicity 2) with corresponding eigenvectors $(1, -1, -1), (1, -1, 1), (1, 0, 0)$. The matrices K, L, M, Q have the same eigenvalues as J . The matrix M has eigenvalues $1, i, -i$ and corresponding eigenvectors $(1, 0, 0), (1, -1, -i), (1, -1, i)$. The other matrices are similar to those already computed.

We can apply this work to find eigenvalues and eigenvectors for the unitary matrices in $\mathcal{U}[\mathcal{G}(\mathcal{S}_3)]$. For example, we have found the eigenvalues $\lambda_1, \lambda_2, \lambda_3$ and corresponding eigenvectors u_1, u_2, u_3 for $A \in \mathcal{G}(\mathcal{S}_3)$. Now $\mathcal{U}(A) = U^{-1}AU$ so the eigenvalues of $\mathcal{U}(A)$ are $\lambda_j, j = 1, 2, 3$ with corresponding eigenvectors $U^{-1}u_j, j = 1, 2, 3$. In particular, the eigenvalues of $\mathcal{U}(A)$ are $1, e^{2\pi i/3}, e^{-2\pi i/3}$ with corresponding eigenvectors

$$U^{-1} = \begin{bmatrix} 1 \\ 1 \\ 1 \end{bmatrix} = \begin{bmatrix} 2 \\ 2\sqrt{3} \\ \sqrt{2/3} \end{bmatrix}$$

$$U^{-1} = \begin{bmatrix} e^{2\pi i/3} \\ 1 \\ e^{-2\pi i/3} \end{bmatrix} = \frac{1}{4} \begin{bmatrix} 5 + i\sqrt{3} \\ 5\sqrt{3} - i \\ 2^{3/2}(1/\sqrt{3} - i) \end{bmatrix}$$

$$U^{-1} = \begin{bmatrix} e^{-2\pi i/3} \\ 1 \\ e^{2\pi i/3} \end{bmatrix} = \frac{1}{4} \begin{bmatrix} 5 - i\sqrt{3} \\ 5\sqrt{3} + i \\ 2^{3/2}(1/\sqrt{3} + i) \end{bmatrix}$$

It is useful to find the eigenvalues and eigenvectors for the unitary operators $\mathcal{U}[\mathcal{G}(\mathcal{S}_3)]$. This is because the self-adjoint generator of $\mathcal{U}(Z), Z \in \mathcal{G}(\mathcal{S}_3)$, which gives an angular momentum operator can then be derived. For example

if $\mathcal{U}(A) = e^{iA'}$ has generator A' , then the eigenvalues of A' are $0, -2\pi/3, 2\pi/3$ and the eigenvectors of A' are the same as those for $\mathcal{U}(A)$.

We now construct a unitary representation of $\mathcal{U}[\mathcal{G}(\mathcal{S}_3)]$ on the 2-dimensional Hilbert space \mathbb{C}^2 . The standard construction goes as follows [7, 8]. For $\widehat{A} \in \mathcal{U}[\mathcal{G}(\mathcal{S}_3)]$, the corresponding unitary matrix on \mathbb{C}^2 is given by

$$\mathcal{R}'(\widehat{A}) = \begin{bmatrix} a & b \\ -\bar{b} & \bar{a} \end{bmatrix}$$

and $a, b \in \mathbb{C}$ satisfy the following seven equations. In these equations, we denote a vector $u \in \mathbb{R}^3$ by its Cartesian coordinates $u = (u_1, u_2, u_3)$.

$$a^2 - b^2 = (\widehat{A}\widehat{e}_1)_1 - i(\widehat{A}\widehat{e}_1)_2 \quad (3.1)$$

$$ab = -\frac{1}{2}(\widehat{A}\widehat{e}_3)_1 + \frac{i}{2}(\widehat{A}\widehat{e}_3)_2 \quad (3.2)$$

$$a\bar{b} = \frac{1}{2}(\widehat{A}\widehat{e}_1)_3 + \frac{i}{2}(\widehat{A}\widehat{e}_2)_2 \quad (3.3)$$

$$\operatorname{Re}(a^2 + b^2) = (\widehat{A}\widehat{e}_2)_2 \quad (3.4)$$

$$\operatorname{Im}(b^2 - \bar{a}^2) = (\widehat{A}\widehat{e}_2)_1 \quad (3.5)$$

$$|b|^2 = \frac{1}{2} \left[1 - (\widehat{A}\widehat{e}_3)_3 \right] \quad (3.6)$$

$$|a|^2 + |b|^2 = 1 \quad (3.7)$$

We define the unitary representation \mathcal{R} of $\mathcal{G}(\mathcal{S}_3)$ on \mathbb{C}^2 by $\mathcal{R}(Z) = \mathcal{R}'[\mathcal{U}(X)]$, $Z \in \mathcal{G}(\mathcal{S}_3)$. Following the above procedure we obtain

$$\mathcal{R}(M) = i \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix}$$

We also have that

$$\mathcal{R}(N) = \frac{1}{\sqrt{3}} \begin{bmatrix} \sqrt{2}e^{-i\pi/6} & e^{i\pi/3} \\ -e^{-i\pi/3} & \sqrt{2}e^{i\pi/6} \end{bmatrix}$$

Since M and N generate $\mathcal{G}(\mathcal{S}_3)$ we can obtain $\mathcal{R}(Z)$ for every $Z \in \mathcal{G}(\mathcal{S}_3)$ by repeated applications of $\mathcal{R}(M)$ and $\mathcal{R}(N)$ using Table 1. These are now listed.

$$\mathcal{R}(A) = \frac{i}{\sqrt{3}} \begin{bmatrix} e^{i\pi/3} & \sqrt{2}e^{-i\pi/6} \\ \sqrt{2}e^{i\pi/6} & -e^{-i\pi/3} \end{bmatrix}, \quad \mathcal{R}(B) = -\frac{1}{\sqrt{3}} \begin{bmatrix} e^{i\pi/6} & \sqrt{2}e^{i\pi/3} \\ -\sqrt{2}e^{-i\pi/3} & e^{-i\pi/6} \end{bmatrix}$$

$$\begin{aligned}
\mathcal{R}(C) &= \frac{1}{\sqrt{3}} \begin{bmatrix} e^{i5\pi/6} & -\sqrt{2} \\ \sqrt{2} & e^{-i\pi/6} \end{bmatrix}, & \mathcal{R}(D) &= -\frac{1}{\sqrt{3}} \begin{bmatrix} e^{i\pi/6} & -\sqrt{2} \\ \sqrt{2} & e^{-i\pi/6} \end{bmatrix} \\
\mathcal{R}(E) &= \frac{i}{\sqrt{3}} \begin{bmatrix} -e^{-i\pi/3} & \sqrt{2}e^{i\pi/6} \\ \sqrt{2}e^{-i\pi/6} & e^{i\pi/3} \end{bmatrix}, & \mathcal{R}(F) &= \frac{1}{\sqrt{3}} \begin{bmatrix} e^{i5\pi/6} & \sqrt{2}e^{i2\pi/3} \\ -\sqrt{2}e^{-i2\pi/3} & e^{-i5\pi/6} \end{bmatrix} \\
\mathcal{R}(G) &= i \begin{bmatrix} e^{i\pi/6} & 0 \\ 0 & -e^{-i\pi/6} \end{bmatrix}, & \mathcal{R}(H) &= i \begin{bmatrix} -e^{-i\pi/6} & 0 \\ 0 & -e^{i\pi/6} \end{bmatrix} \\
\mathcal{R}(I) &= \begin{bmatrix} 1 & 0 \\ 0 & 1 \end{bmatrix}, & \mathcal{R}(J) &= \frac{1}{\sqrt{3}} \begin{bmatrix} -i & -\sqrt{2} \\ \sqrt{2} & i \end{bmatrix} \\
\mathcal{R}(K) &= \frac{1}{\sqrt{3}} \begin{bmatrix} e^{-i\pi/2} & \sqrt{2}e^{i\pi/3} \\ -\sqrt{2}e^{-i\pi/3} & e^{i\pi/2} \end{bmatrix}, & \mathcal{R}(L) &= -\frac{1}{\sqrt{3}} \begin{bmatrix} e^{i\pi/2} & \sqrt{2}e^{i2\pi/3} \\ -\sqrt{2}e^{-i2\pi/3} & e^{-i\pi/2} \end{bmatrix} \\
\mathcal{R}(O) &= \frac{1}{\sqrt{3}} \begin{bmatrix} \sqrt{2}e^{-i\pi/6} & e^{-i\pi/3} \\ -e^{i\pi/3} & \sqrt{2}e^{i\pi/6} \end{bmatrix}, & \mathcal{R}(P) &= \frac{i}{\sqrt{3}} \begin{bmatrix} \sqrt{2} & e^{-i\pi/6} \\ e^{i\pi/6} & -\sqrt{2} \end{bmatrix} \\
\mathcal{R}(Q) &= \frac{i}{\sqrt{3}} \begin{bmatrix} \sqrt{2} & -e^{-i\pi/6} \\ -e^{i\pi/6} & -\sqrt{2} \end{bmatrix}, & \mathcal{R}(R) &= -\frac{1}{\sqrt{3}} \begin{bmatrix} \sqrt{2}e^{i\pi/6} & -e^{-i\pi/3} \\ e^{i\pi/3} & \sqrt{2}e^{-\pi/6} \end{bmatrix} \\
\mathcal{R}(S) &= \frac{1}{\sqrt{3}} \begin{bmatrix} \sqrt{2}e^{i\pi/6} & -e^{i\pi/3} \\ -e^{i\pi/6} & -\sqrt{2} \end{bmatrix}, & \mathcal{R}(T) &= \begin{bmatrix} 0 & -e^{i\pi/6} \\ e^{-i\pi/6} & 0 \end{bmatrix} \\
\mathcal{R}(U) &= \begin{bmatrix} 0 & -e^{-i\pi/6} \\ e^{i\pi/6} & 0 \end{bmatrix}, & \mathcal{R}(V) &= \frac{1}{\sqrt{3}} \begin{bmatrix} i\sqrt{2} & 1 \\ -1 & -i\sqrt{2} \end{bmatrix} \\
\mathcal{R}(W) &= \frac{1}{\sqrt{3}} \begin{bmatrix} \sqrt{2}e^{i5\pi/6} & 1 \\ -1 & \sqrt{2}e^{-i5\pi/6} \end{bmatrix}, & \mathcal{R}(X) &= \frac{i}{\sqrt{3}} \begin{bmatrix} \sqrt{2}e^{i\pi/6} & 1 \\ -1 & \sqrt{2}e^{-i\pi/6} \end{bmatrix}
\end{aligned}$$

Technically speaking, \mathcal{R} is a *projective representation* of $\mathcal{G}(\mathcal{S}_3)$ on \mathbb{C}^2 .

That is $\mathcal{R}(YZ) = \pm\mathcal{R}(Y)\mathcal{R}(Z)$ in general. However, this is not important because quantum states are only determined within a scalar multiple of absolute value one. Examples are $GH = I$ but $\mathcal{R}(G)\mathcal{R}(H) = -I$ and $J^2 = I$ and $\mathcal{R}(J)^2 = -I$.

4 Discrete Spacetime

We have previously considered the space lattice \mathcal{S}_3 . We now adjoin time to obtain a spacetime lattice. Let $d, e, f, g \in \mathbb{R}$ be unit vectors satisfying

$$\langle d, e \rangle = \langle d, f \rangle = \langle d, g \rangle = 0$$

and

$$\langle e, f \rangle = \langle e, g \rangle = \langle f, g \rangle = 1/2$$

Then

$$\mathcal{S}_4 = \{td + ne + pf + qg : t, n, p, q \in \mathbb{Z}\}$$

is a 4-dimensional module over \mathbb{Z} with basis $\{d, e, f, g\}$. Clearly, \mathcal{S}_3 is a 3-dimensional submodule of \mathcal{S}_4 . We are mainly concerned with the subset $\mathcal{S}_4^+ \subseteq \mathcal{S}_4$ of vectors with $t \geq 0$. We frequently call these vectors *vertices* and consider them to be tiny spacetime cells that may be occupied by a particle. If $u = td + ne + pf + qg$, we write $u = (t, n, p, q)$ and use the notation $u^0 = t$, $u^1 = n$, $u^2 = p$, $u^3 = q$.

We define the usual norm $\|\cdot\|_3$ on \mathcal{S}_3 by

$$\|ne + pf + qg\|_3^2 = n^2 + p^2 + q^2 + np + nq + pq$$

and the indefinite norm $\|\cdot\|_4$ on \mathcal{S}_4 by

$$\|td + ne + pf + qg\|_4^2 = t^2 - \|ne + pf + qg\|_3^2$$

We sometimes write $\|u\|_4 = (u^0)^2 - \|\underline{u}\|_3^2$ where $\underline{u} = (u^1, u^2, u^3) \in \mathbb{R}^3$. Of course, the *spacetime distance between* $u, v \in \mathcal{S}_4$ is $\|u - v\|_4$. For $t \geq 0$, we define

$$\mathcal{S}^t = \{u \in \mathcal{S}_4^+ : u^0 = t, \|u\|_4^2 \geq 0\}$$

and we call \mathcal{S}^t the *universe at time* t . We call

$$\overline{\mathcal{S}}^t = \cup \{\mathcal{S}^{t'} : 0 \leq t' \leq t, t' \in \mathbb{Z}\}$$

the *universe history until time t*. For $u \in \mathcal{S}_4^+$, the *forward light cone at u* is

$$\mathcal{C}_u^+ = \{v \in \mathcal{S}_4^+ : v^0 \geq u^0, \|v - u\|_4^2 \geq 0\}$$

As usual \mathcal{C}_u^+ is the set of vertices that u can reach with a physical signal. In other words \mathcal{C}_u^+ is the set of vertices that u can influence. Of course,

$$\mathcal{C}_0^+ = \{u \in \mathcal{S}_4 : u^0 \geq 0, \|u\|_4 \geq 0\} = \bigcup_{t \geq 0} \mathcal{S}^t = \bigcup_{t \geq 0} \overline{\mathcal{S}}^t$$

We consider \mathcal{C}_0^+ to be the spacetime background and shall only consider vertices in \mathcal{C}_0^+ . The *forward null surface at u* $u \in \mathcal{C}_0^+$ is

$$\eta_u^+ = \{v \in \mathcal{S}_4 : v^0 \geq u^0, \|v - u\|_4^2 = 0\}$$

We interpret η_u^+ as the vertices that u can reach with a light signal. For $u, v \in \mathcal{C}_0^+$ we write $u < v$ and say that v is in the *causal future* of u if $u^0 < v^0$ and $\|v - u\|_4^2 \geq 0$. Of course, $u < v$ if and only if $v \in \mathcal{C}_u^+ \setminus \{u\}$.

Theorem 4.1. *The relation $<$ is a partial order on \mathcal{C}_0^+ .*

Proof. Clearly $u \not< u$ for all $u \in \mathcal{C}_0^+$. Suppose $u, v, w \in \mathcal{C}_0^+$ with $u < v$ and $v < w$. Then $u^0 < v^0 < w^0$ and

$$\begin{aligned} (v^0 - u^0)^2 - \|\underline{v} - \underline{u}\|_3^2 &= \|v - u\|_4^2 \geq 0 \\ (w^0 - v^0)^2 - \|\underline{w} - \underline{v}\|_3^2 &= \|w - v\|_4^2 \geq 0 \end{aligned} \tag{4.1}$$

By the triangle inequality we have

$$\begin{aligned} \|w - u\|_4^2 &= (w^0 - u^0)^2 - \|\underline{w} - \underline{u}\|_3^2 \geq (w^0 - u^0)^2 - \|\underline{w} - \underline{v}\|_3^2 - \|\underline{v} - \underline{u}\|_3^2 \\ &= (w^0 - v^0)^2 + (v^0 - u^0)^2 + 2(w^0 - v^0)(v^0 - u^0) - \|\underline{w} - \underline{v}\|_3^2 - \|\underline{v} - \underline{u}\|_3^2 \end{aligned}$$

Then $\|w - u\|_4^2 \geq 0$ follows from (4.1). \square

The partial order $<$ restricted to $\overline{\mathcal{S}}^t$ makes $\overline{\mathcal{S}}^t$ a finite poset that is frequently called a *causal set* or *causet* [2–4]. If $u < v$ and there is no $w \in \mathcal{C}_0^+$ with $u < w < v$, we say that v is a *child* of u and u is a *parent* of v and write $u \prec v$. If $u \prec v$, we call the edge uv a *link from u to v*. A *path from u to v* is a sequence $w_1 \prec w_2 \prec \dots \prec w_n$ where $w_1 = u$ and $w_n = v$. We call $n - 1$ the *length* of this path. In general, there may not be a path from u to v .

Theorem 4.2. (i) *There exists a path from u to v if and only if $u < v$.*
(ii) *$u \prec v$ if and only if $v^0 = u^0 + 1$ and $\|v - u\|_4 = 0$ or 1.* (iii) *If $u < v$, then any two paths from u to v have length $v^0 - u^0$.*

Proof. (i) If there exists a path from u to v , then $u < v$ by transitivity. Conversely, suppose $u < v$. If $u \prec v$, the u, v form a path and we are finished. If $u \not\prec v$, then there is a $w \in \mathcal{C}_0^+$ such that $u < w < v$. If $u \prec w \prec v$, then u, w, v form a path. Otherwise, there is a $w_1 \in \mathcal{C}_0^+$ such that $u < w_1 < w < v$ or $u < w < w_1 < v$. Since there are only a finite number of vertices between u and v , this process must eventually end and we obtain a path from u to v . (ii) In Lemma 2.1 we constructed the 12 unit vectors in \mathcal{S}_3 . It will be convenient to label them by $\underline{e}_1 = e, \underline{e}_2 = f, \underline{e}_3 = g, \underline{e}_4 = e - f, \underline{e}_5 = e - g, \underline{e}_6 = f - g, \underline{e}_7 = -e, \underline{e}_8 = -f, \dots, \underline{e}_{12} = g - f$. We can construct \mathcal{C}_0^+ as follows. The universe histories $\overline{\mathcal{S}}^0$ and $\overline{\mathcal{S}}^1$ are given by $\overline{\mathcal{S}}^0 = \{0\}$,

$$\overline{\mathcal{S}}^1 = \{0, (1, \underline{0}), (1, \underline{e}_1), (1, \underline{e}_2), \dots, (1, \underline{e}_{12})\}$$

To obtain $\overline{\mathcal{S}}^2$ we add the vectors $(1, \underline{0}), (1, \underline{e}_1), \dots, (1, \underline{e}_{12})$ to those in $\overline{\mathcal{S}}^1$

$$\begin{aligned} \overline{\mathcal{S}}^2 = \{ & 0, (1, \underline{0}), (1, \underline{e}_1), \dots, (1, \underline{e}_{12}), (2, \underline{0}), (2, \underline{e}_1), (2, 2\underline{e}_1), (2, \underline{e}_1 + \underline{e}_2), \\ & \dots, (2, \underline{e}_{11} + \underline{e}_{12}) \} \end{aligned}$$

Strictly speaking, there are repeats in this list for $\overline{\mathcal{S}}^2$ which should be eliminated. Continue this process to obtain the universe histories $\overline{\mathcal{S}}^t$ for all $t \in \mathbb{N}$. If $v^0 = u^0 + 1$ and $\|v - u\|_4 = 0$ or 1, then $u < v$. Suppose there is a $w \in \mathcal{C}_0^+$ such that $u < w < v$. Then $u^0 < w^0 < v^0$ which is impossible so $u \prec v$. Conversely, suppose $u \prec v$. By our previous construction of \mathcal{C}_0^+ we have that

$$(v^0, \underline{v}) = (u^0, \underline{u}) + (1, \underline{0}) = (u^0 + 1, \underline{u})$$

or for some $j = 1, 2, \dots, 12$

$$(v^0, \underline{v}) = (u^0, \underline{u}) + (1, \underline{e}_j) = (u^0 + 1, \underline{u} + \underline{e}_j)$$

In either case, $v^0 = u^0 + 1$ and in the first case

$$\|v - u\|_4^2 = 1 - 0 = 1$$

while in the second case

$$\|v - u\|_4^2 = 1 - \|\underline{e}_j\|_3^2 = 1 - 1 = 0$$

(iii) If $u < v$ then $u \in \mathcal{S}^t$ and $v \in \mathcal{S}^{t'}$ for $t < t'$. From (i) there is a path from u to v and by the construction in (ii), it is clear that every path from u to v has length $t' - t$. \square

Applying Theorem 4.2(iii), all paths from 0 to v have length v^0 . We then say that the *height* of v is v^0 . We conclude that \mathcal{S}^t is precisely the set of vertices with height t and call \mathcal{S}^t the *t-th shell* in $\overline{\mathcal{S}}^{t'}$ for $t \leq t'$. Also, notice from the construction in Theorem 4.2(ii) that every $u \in \mathcal{C}_0^+$ has precisely 13 children. We do not know the number of parents a vertex has and it would be interesting to find out. This would be useful in finding the cardinalities of \mathcal{S}^t and $\overline{\mathcal{S}}^t$ which are also unknown.

If a causet P has the property that whenever $u, v \in P$ with $u < v$, then any two paths from u to v have the same length, we call P *weakly covariant*. For a general causet P the *height* $h(v)$ of $v \in P$ is the length of the longest path terminating at v . We say that P is *covariant* if $h(u) < h(v)$ implies that $u < v$ [2–4]. We have shown in previous works that a covariant causet has a natural weak metric and notions of curvature and geodesics. We have employed these concepts to develop a discrete quantum gravity [2–4].

Lemma 4.3. *A covariant causet is weakly covariant.*

Proof. Let P be a covariant causet and let $v \in P$ with $v > 0$. Let $w_1 \prec w_2 \prec \dots \prec w_n$ be a path from 0 to v . It follows from covariance that $h(w_j) = j$ so that $h(v) = h(w_n) = n$. We conclude that any path from 0 to v has length n . Moreover, if $u < v$, then any path from u to v has length $h(v) - h(u)$. \square

Applying Theorem 4.2(iii) we conclude that the causets $\overline{\mathcal{S}}^t$, $t > 2$, are weakly covariant and it is easy to check that they are not covariant so the converse of Lemma 4.3 does not hold. However, $\overline{\mathcal{S}}^t$ for t up to some limit t_1 (possibly about 300) describes an inflationary period in which the universe histories are essentially flat and gravity has not yet taken effect. After time t_1 most of the spacetime cells have been formed and the system goes into a multiverse period. During this period, there are myriads of possible universes which are expanding much more slowly. The universes develop curvatures which are the cause of gravity. In the multiverse period, the universe histories become covariant and this is employed to describe curvature [3]. However, the curvature is local and the universes are essentially flat because most of the spacetime cells have already been formed.

We assume that if a particle moves from u to v , $u, v \in \mathcal{C}_0^+$ with $u < v$, then it traverses a path $w_1 \prec w_2 \prec \dots \prec w_n$, from u to v where $w_1 = u$, $w_n = v$. The length of the path is $n - 1 = v^0 - u^0$. By Theorem 4.2(ii) the time interval between w_j and w_{j+1} is 1 unit and the space distance moved is either 0 or 1 unit. This tells us that the *instantaneous speed* of a particle is either 0 or 1. Thus, a particle that is not motionless can move at only one speed. If we are indeed using Planck units, this speed is the speed of light c in a vacuum. This fundamental principle is the reason that c is the speed limit for physical signals. What about objects that we know move slower than c ? They appear to be moving slower than c because we are actually measuring average speeds in our observations. If a particle propagates from u to v , we define its *average speed* to be $s_a = \|\underline{v} - \underline{u}\|_3 / (v^0 - u^0)$. Of course, $0 \leq s_a \leq 1$ so c is still the speed limit for average speed. Notice that s_a can have various values. For example, if $u = 0$ and $v = 3d + 3e$, then $s_a = 1$, while if $u = 0$ and $v = 3d + 2e - f$ then $s_a = 1/\sqrt{3}$.

It is interesting to find the possible average speeds. Of course, there are only a finite number of possibilities for $u, v \in \overline{\mathcal{S}}^t$ with t fixed. For simplicity, suppose a particle propagates from 0 to v so that $s_a = \|\underline{v}\|_3 / v^0$. For $v^0 = 1$, the only possibilities are $s_a = 0, 1$. For $v^0 = 2$, the possibilities are $s_a = 0, 1/2, \sqrt{2}/2, \sqrt{3}/2, 1$. For $v^0 = 3$, we have

$$s_a = 0, 1/3, \sqrt{2}/3, \sqrt{3}/3, \dots, \sqrt{8}/\sqrt{3}, 1$$

But now the pattern ends and there are gaps. For $v^0 = 4$ we have

$$s_a = 0, \sqrt{3}/4, \sqrt{4}/4, \sqrt{5}/4, \dots, \sqrt{15}/4, 1$$

For $v^0 = 5$ we obtain

$$s_a = 0, 1/5, \sqrt{2}/5, 2/5, \sqrt{6}/5, \sqrt{7}/5, 3/5, \sqrt{12}/5, \sqrt{13}/5, \dots, \sqrt{24}/5, 1$$

We do not know a general formula for possible s_a corresponding to v^0 .

A *symmetry* on \mathcal{S}_4 is a linear bijection $T: \mathcal{S}_4 \rightarrow \mathcal{S}_4$ that preserves the norm $\|\cdot\|^4$ and has unit determinant. A symmetry T is a *boost* if $Td \neq d$. Notice, if $Td = d$ then $T\mathcal{S}_3 = \mathcal{S}_3$ because

$$0 = \langle d, e \rangle = \langle Td, Te \rangle = \langle d, Te \rangle$$

so $Te \in \mathcal{S}_3$ and similarly, $Tf, Tg \in \mathcal{S}_3$. Symmetries that are not boosts have the form $1 \oplus Z$, $Z \in \mathcal{G}(\mathcal{S}_3)$. The question that now presents itself is: Are

there any boosts? The simplest examples in classical special relativity are the *basic* boosts of the form [7, 8]

$$T = \begin{bmatrix} 1/\sqrt{1-v^2} & v/\sqrt{1-v^2} & 0 & 0 \\ v/\sqrt{1-v^2} & 1/\sqrt{1-v^2} & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

The transformation T describes a coordinate system moving at the constant velocity v along the x^1 axis. But in our situation, we only have two values for an instantaneous velocity, $v = 0$ or 1 . In the first case, $T = I$ which is not a boost, while the second case is impossible. Even if other values of v are allowed (average velocities), it is easy to check that T does not leave \mathcal{S}_4 invariant.

Theorem 4.4. *There are no boosts.*

Proof. Suppose $T: \mathcal{S}_4 \rightarrow \mathcal{S}_4$ is a boost and

$$\begin{aligned} Td &= t_0d + t_1e + t_2f + t_3g \\ Te &= s_0d + s_1e + s_2f + s_3g \end{aligned}$$

Then

$$1 = \|d\|_4^2 = \|Td\|_4^2 = t_0^2 - \frac{1}{2} [(t_1 + t_2)^2 + (t_1 + t_3)^2 + (t_2 + t_3)^2]$$

Hence,

$$(t_1 + t_2)^2 + (t_1 + t_3)^2 + (t_2 + t_3)^2 + 2 = 2t_0^2 \quad (4.2)$$

The only nontrivial solutions to (4.2) are

$$t_0 = 2, t_1 = t_2 = 1, t_3 = 0 \quad (4.3)$$

$$t_0 = 3, t_1 = t_2 = 2, t_3 = -2 \quad (4.4)$$

and similar ones. Notice that trivial solutions like $t_0 = 1, t_1 = t_2 = t_3 = 0$ do not give boosts because then $Td = d$. We also have

$$-1 = \|e\|_4^2 = \|Te\|_4^2 = s_0^2 - \frac{1}{2} [(s_1 + s_2)^2 + (s_1 + s_3)^2 + (s_2 + s_3)^2]$$

Hence,

$$(s_1 + s_2)^2 + (s_1 + s_3)^2 + (s_2 + s_3)^2 = 2s_0^2 + 2 \quad (4.5)$$

The only nontrivial solutions of (4.5) are

$$s_0 = 1, \quad s_1 = s_2 = 1, \quad s_3 = -1 \quad (4.6)$$

$$s_0 = 2, \quad s_1 = 2, \quad s_2 = 1, \quad s_3 = -1 \quad (4.7)$$

and similar ones. We also must satisfy

$$0 = \langle d, e \rangle = \langle Td, Te \rangle = t_0 s_0 - \frac{1}{2}(t_1 s_1 + t_2 s_2 + t_3 s_3) \quad (4.8)$$

The only combination of cases (4.3), (4.4) with cases (4.6), (4.7) that satisfy (4.8) are (4.4) and (4.6). Thus,

$$Te = d + e + f - g$$

But the same reasoning shows that $Tf = d + e + f - g$. Hence, $Te = Tf$ which contradicts the bijectivity of T . \square

We conclude from Theorem 4.4 that all symmetries of \mathcal{S}_4 have the form $1 \oplus Z$, $Z \in \mathcal{G}(\mathcal{S}_3)$. We denote this group by $\mathcal{G}(\mathcal{S}_4)$.

5 Discrete Quantum Field Theory

As discussed in the previous section, \mathcal{S}_4 is the discrete spacetime background for expanding universe histories during an inflationary period. Most of the spacetime cells are formed and the universe is essentially flat during this period. Curvatures and gravity do not emerge until later. We now discuss discrete quantum field theory. As usual, this theory only employs special relativity and the general relativity of gravity is neglected. This assumption is well-founded because gravitation is extremely weak compared to electromagnetic and nuclear forces. For this reason we shall take \mathcal{S}_4 as the underlying spacetime for our quantum field theory.

Let $\widehat{\mathcal{S}}_4$ be a copy of \mathcal{S}_4 whose elements are labeled by

$$p = (p^0, \underline{p}) = (p^0, p^1, p^2, p^3)$$

in Cartesian coordinates. We think of $\widehat{\mathcal{S}}_4$ as being dual to \mathcal{S}_4 with indefinite inner product

$$p \cdot x = p^0 x^0 - p^1 x^1 - p^2 x^2 - p^3 x^3$$

where $x \in \mathcal{S}_4$ is given in Cartesian coordinates. We call p^0 the *total energy*, \underline{p} the *momentum* and

$$m^2 = \|\underline{p}\|_4^2 = (p^0)^2 - \|\underline{p}\|_3^2 = (p^0)^2 - (p^1)^2 - (p^2)^2 - (p^3)^2$$

the *mass squared*. We only consider m^2 when $m^2 \geq 0$. We see that m^2 can only have integer values so mass is discrete in this theory. We have not found a formula for the possible values of m^2 as a function of p^0 and this would be of interest to know. We have computed the values of $\|\underline{p}\|_3^2$ up to 49 which are the following:

$$\|\underline{p}\|_3^2 = 0, 1, 2, \dots, 13, 16, 18, 19, 21, 23, \dots, 28, 31, 33, 35, \dots, 39, 43, 49$$

We then obtain the following mass squared values as a function of p^0

p^0	m^2
0	0
1	0, 1
2	0, 1, 2, 3, 4
3	0, 1, 2, ..., 9
4	0, 3, 4, ..., 16
5	0, 1, 2, 4, 6, 7, 9, 12, 13, ..., 25
6	0, 1, 3, 5, 8, 9, ..., 13, 15, 17, 18, 20, 23, 24, ..., 36
7	0, 6, 10, ..., 14, 16, 18, 21, ..., 26, 28, 30, 31, 33, 36, ..., 49

We can identify $\mathcal{G}(\mathcal{S}_4)$ with $\mathcal{G}(\widehat{\mathcal{S}}_4)$ and when we write $Z \in \mathcal{G}(\mathcal{S}_4)$ we mean $Z \in \mathcal{G}(\mathcal{S}_4)$ or $Z \in \mathcal{G}(\widehat{\mathcal{S}}_4)$ whichever is applicable. The set $\mathcal{S}_4 \times \mathcal{G}(\mathcal{S}_4)$ becomes a group with product

$$(y, Y)(z, Z) = (y + Yz, YZ)$$

We think of $\mathcal{S}_4 \times \mathcal{G}(\mathcal{S}_4)$ as a discrete Poincare group. For $m^2 \geq 0$, the *mass hyperboloid* is the set

$$\Gamma_m = \left\{ p \in \widehat{\mathcal{S}}_4 : \|\underline{p}\|_4^2 = m^2 \right\}$$

The basic Hilbert space for this theory is

$$H_m = L_2(\Gamma_m) = \left\{ f : \Gamma_m \rightarrow \mathbb{C} : \sum_{p \in \Gamma_m} |f(p)|^2 < \infty \right\}$$

with the usual inner product

$$\langle f, g \rangle = \sum_{p \in \Gamma_m} \overline{f(p)} g(p)$$

Define the representation V of $\mathcal{S}_4 \times \mathcal{G}(\mathcal{S}_4)$ on H_m by

$$[V(y, Y)f](p) = e^{ip \cdot y} f(Y^{-1}p)$$

To show that V gives a representation, we have

$$\begin{aligned} [V(y, Y)V(z, Z)f](p) &= e^{ip \cdot y} [V(z, Z)f](Y^{-1}p) \\ &= e^{ip \cdot y} e^{iY^{-1}p \cdot z} f(Z^{-1}Y^{-1}p) \\ &= e^{ip \cdot (y + Yz)} f((YZ)^{-1}p) \\ &= [V(y + Yz, YZ)f](p) \\ &= [V((y, Y)(z, Z))f](p) \end{aligned}$$

where the third equality follows from

$$Y^{-1}p \cdot z = YY^{-1}p \cdot Yz = p \cdot Yz$$

Notice that $Z \in \mathcal{G}(\mathcal{S}_4)$ leaves Γ_m invariant because

$$\|Zp\|_4^2 = \|p\|_4^2 = m^2$$

It follows that V is a unitary representation. Indeed,

$$\begin{aligned} \langle V(z, Z)f, V(z, Z)g \rangle &= \sum_{p \in \Gamma_m} \overline{[V(z, Z)f](p)} [V(z, Z)g](p) \\ &= \sum_{p \in \Gamma_m} e^{-ip \cdot z} \overline{f(Z^{-1}p)} e^{ip \cdot z} g(Z^{-1}p) \\ &= \sum_{p \in \Gamma_m} \overline{f(p)} g(p) = \langle f, g \rangle \end{aligned}$$

We have the four self-adjoint operators P^j , $j = 0, 1, 2, 3$, on H_m given by $(P^j f)(p) = p^j f(p)$. Since

$$\begin{aligned} (P^0)^2 f(p) &= (p^0)^2 f(p) = [(p^1)^2 + (p^2)^2 + (p^3)^2 + m^2] f(p) \\ &= [(P^1)^2 + (P^2)^2 + (P^3)^2 + m^2] f(p) \end{aligned}$$

we conclude that

$$(P^0)^2 = (P^1)^2 + (P^2)^2 + (P^3)^2 + m^2 I$$

The eigenvectors of P^2 are the characteristic functions χ_p , $p \in \Gamma_m$ with eigenvalues p^j , $j = 0, 1, 2, 3$, on the mass hyperboloid.

The space H_m that we have considered until now is the *scalar* (or *spin-0*) *mass m* Hilbert space. The *vector* (or *spin-1*) *mass m* Hilbert space is $H_m \otimes \mathbb{C}^3$ with the usual inner product. The representation V_1 of $\mathcal{S}_4 \times \mathcal{G}(\mathcal{S}_4)$ on $H_m \otimes \mathbb{C}^3$ is given by

$$V_1(y, Y)f(p) \otimes v = e^{ip \cdot y} f(Y^{-1}p) \otimes U(Y^{-1})v$$

As before, V_1 is a unitary representation. The *spin-1/2 mass m* Hilbert space is $H_m \otimes \mathbb{C}^2$ with the usual inner product. The unitary representation $V_{1/2}$ of $\mathcal{S}_4 \times \mathcal{G}(\mathcal{S}_4)$ on $H_m \otimes \mathbb{C}^2$ is given by

$$V_{1/2}(y, Y)f(p) \otimes v = e^{ip \cdot y} f(Y^{-1}p) \otimes \mathcal{R}(Y^{-1})v$$

We can continue to form the *spin-n/2 mass m* Hilbert space

$$H_m \otimes \mathbb{C}^2 \otimes \dots \otimes \mathbb{C}^2$$

where there are n factors of \mathbb{C}^2 with the unitary representation

$$V_{n/2}(y, Y)f(p) \otimes v_1 \otimes \dots \otimes v_n = e^{ip \cdot y} f(Y^{-1}p) \otimes \mathcal{R}(Y^{-1})v_1 \otimes \dots \otimes \mathcal{R}(Y^{-1})v_n$$

We finally develop a discrete quantum field theory. For simplicity, we shall only consider scalar fields and the extension to nonzero spins is fairly straightforward. We first form the Hilbert space $\mathcal{H}_m = \bigoplus_{n=0}^{\infty} H_m^n$ where $H_m^0 = \mathbb{C}$ and for $n \in \mathbb{N}$, H_m^n is the symmetric tensor product

$$H_m^n = H_m \mathbb{S} H_m \mathbb{S} \dots \mathbb{S} H_m \quad (n \text{ factors})$$

We can consider $f \in H_m^n$ as a symmetric function of n variables $f(p_1, \dots, p_n)$, $p_j \in \Gamma_m$, $j = 1, 2, \dots, n$. For $x \in \mathcal{S}_4$ we construct the *field operators* $\phi(x): \mathcal{H}_m \rightarrow \mathcal{H}_m$ by defining $\phi(x): H_m^{n+1} \rightarrow H_m^n$ as follows. For $f \in H_m^n$ we have

$$f = f^{(0)} \oplus f^{(1)} \oplus f^{(2)} \oplus \dots \quad (5.1)$$

where $f^j \in H_m^j$, $j = 0, 1, 2, \dots$. We define $\phi(x)f^{(0)} = 0$ and for $n \in \mathbb{N}$ we define [7, 8]

$$[\phi(x)f^{(n+1)}](p_1, p_2, \dots, p_n) = \sqrt{n+1} \sum_{p \in \Gamma_m} e^{-ip \cdot x} f^{(n+1)}(p, p_1, \dots, p_n)$$

We also construct the field operators $\psi(x): \mathcal{H}_m \rightarrow \mathcal{H}_m$ by defining $\psi(x): H_m^n \rightarrow H_m^{n+1}$ as follows

$$[\psi(x)f^{(n)}](p_1, p_2, \dots, p_{n+1}) = \frac{1}{\sqrt{n+1}} \sum_{j=1}^{n+1} e^{ip_j \cdot x} f^{(n)}(p_1, \dots, \widehat{p}_j, \dots, p_{n+1})$$

where \widehat{p}_j means omit p_j . We interpret $\phi(x)$ as an operator that annihilates a mass m particle at the spacetime point x while $\psi(x)$ is an operator that creates a mass m particle at the spacetime point x . We obtain a unitary representation of $\mathcal{S}_4 \times \mathcal{G}(\mathcal{S}_4)$ on \mathcal{H}_m by defining

$$[V(y, Y)f^{(n)}](p_1, \dots, p_n) = e^{i \sum_{j=1}^n p_j \cdot y} f^{(n)}(Y^{-1}p_1, \dots, Y^{-1}p_n)$$

on H_m^n and extending $V(y, Y)$ to \mathcal{H}_m by applying (5.1).

Theorem 5.1. *For every $x \in \mathcal{S}_4$, $\psi(x) = \phi(x)^*$.*

Proof. Let $f, g \in \mathcal{H}_m$. Since f and g are symmetric in their variables and letting $p = p_{n+1}$ in the fourth equality, we obtain

$$\begin{aligned} \langle \phi(x)f^{(n+1)}, g^{(n)} \rangle &= \left\langle \sqrt{n+1} \sum_{p \in \Gamma_m} e^{-ip \cdot x} f^{(n+1)}(p, p_1, \dots, p_n), g^{(n)}(p_1, \dots, p_n) \right\rangle \\ &= \sqrt{n+1} \sum_{p_1, \dots, p_n} \sum_p e^{ip \cdot x} \overline{f^{(n+1)}(p, p_1, \dots, p_n)} g^{(n)}(p_1, \dots, p_n) \\ &= \frac{1}{\sqrt{n+1}} \left\{ \left[\sum_{p, p_1, \dots, p_n} e^{ip \cdot x} \overline{f^{(n+1)}(p, p_1, \dots, p_n)} \right. \right. \\ &\quad + \sum_{p, p_1, \dots, p_n} e^{ip_1 \cdot x} \overline{f^{(n+1)}(p, p_1, \dots, p_n)} + \dots \\ &\quad \left. \left. + \sum_{p, p_1, \dots, p_n} e^{ip_n \cdot x} \overline{f^{(n+1)}(p, p_1, \dots, p_n)} \right] g^{(n)}(p_1, \dots, p_n) \right\} \end{aligned}$$

$$\begin{aligned}
&= \frac{1}{\sqrt{n+1}} \sum_{p_1, \dots, p_n} \overline{f^{(n+1)}(p_1, \dots, p_{n+1})} \\
&\quad \sum_{j=1}^{n+1} e^{ip_j \cdot x} g^{(n)}(p_1, \dots, \widehat{p}_j, \dots, p_{n+1}) \\
&= \frac{1}{\sqrt{n+1}} \left\langle f^{(n+1)}(p_1, \dots, p_{n+1}), \sum_{j=1}^{n+1} e^{ip_j \cdot x} g^{(n)}(p_1, \dots, \widehat{p}_j, \dots, p_{n+1}) \right\rangle \\
&= \langle f^{(n+1)}, \psi(x)g^{(n)} \rangle
\end{aligned}$$

We conclude that

$$\begin{aligned}
\langle \phi(x)f, g \rangle &= \sum_{n=0}^{\infty} \langle \phi(x)f^{(n+1)}, g^{(n)} \rangle = \sum_{n=0}^{\infty} \langle f^{(n+1)}, \psi(x)g^{(n)} \rangle \\
&= \langle f, \psi(x)g \rangle
\end{aligned}$$

Hence, $\psi(x) = \phi(x)^*$. □

It is not hard to show that the commutator $[\phi(x), \phi(y)] = 0$ for all $x, y \in \mathcal{S}_4$. Similarly, $[\psi(x), \psi(y)] = 0$ for all $x, y \in \mathcal{S}_4$.

Theorem 5.2. *For all $x, y \in \mathcal{S}_4$ we have that*

$$[\phi(x), \psi(y)] = \sum_{p \in \Gamma_m} e^{ip \cdot (y-x)} I$$

Proof. For $f^{(n)} \in H_m^n$ we have

$$\begin{aligned}
[\phi(x)\psi(y)f^{(n)}](p_1, \dots, p_n) &= \sqrt{n+1} \sum_{p \in \Gamma_m} e^{-ip \cdot x} [\psi(y)f^{(n)}](p, p_1, \dots, p_n) \\
&= \sum_{p \in \Gamma_m} e^{-ip \cdot x} [e^{ip \cdot y} f^{(n)}(p_1, \dots, p_n) + e^{ip_1 \cdot y} f^{(n)}(p, p_1, \dots, p_n) \\
&\quad + \dots + e^{ip_n \cdot y} f^{(n)}(p, p_1, \dots, p_n)]
\end{aligned}$$

On the other hand,

$$[\psi(y)\phi(x)f^{(n)}](p_1, \dots, p_n) = \frac{1}{\sqrt{n}} \sum_{j=1}^n e^{ip_j \cdot y} [\phi(x)f^{(n)}](p_1, \dots, \widehat{p}_j, \dots, p_n)$$

$$\begin{aligned}
&= \sum_{j=1}^n \sum_{p \in \Gamma_m} e^{-ip \cdot x} e^{ip_j \cdot y} f^{(n)}(p, p_1, \dots, \widehat{p}_j, \dots, p_n) \\
&= \sum_{p \in \Gamma_m} e^{-ip \cdot x} \left[e^{ip_1 \cdot y} f^{(n)}(p, p_2, \dots, p_n) + e^{ip_2 \cdot y} f^{(n)}(p, p_1, p_3, \dots, p_n) \right. \\
&\quad \left. + \dots + e^{ip_n \cdot y} f^{(n)}(p, p_1, p_2, \dots, p_{n+1}) \right]
\end{aligned}$$

Therefore,

$$[\phi(x), \psi(y)] f^{(n)} = \sum_{p \in \Gamma_m} e^{ip \cdot (y-x)} f^{(n)}$$

It follows that

$$[\phi(x), \psi(y)] f = \sum_{p \in \Gamma_m} e^{ip \cdot (y-x)} f$$

for all $f \in \mathcal{H}_m$ □

We also define the self-adjoint field operators $\xi(x) = \phi(x) + \psi(x)$ for all $x \in \mathcal{S}_4$.

Corollary 5.3. *For all $x, y \in \mathcal{S}_4$ we have that*

$$[\xi(x), \xi(y)] = 2i \sum_{p \in \Gamma_m} \sin p \cdot (y - x) I$$

Proof. Since $[\phi(x), \phi(y)] = [\psi(x), \psi(y)] = 0$ we have that

$$[\xi(x), \xi(y)] x = [\phi(x), \psi(y)] + [\psi(x), \phi(y)]$$

Applying Theorem 5.2 gives

$$\begin{aligned}
[\xi(x), \xi(y)] &= \sum_{p \in \Gamma_m} [e^{ip \cdot (y-x)} - e^{-ip \cdot (y-x)}] I \\
&\quad 2i \sum_{p \in \Gamma_m} \sin p \cdot (y - x) I \quad \square
\end{aligned}$$

The quantum field theory that we have developed is essentially trivial because the fields are free with no interactions. We view this as just the beginning, and initiate the difficult task of investigating interacting fields in the next section. It is hoped that these initial steps may result in a mathematically rigorous quantum field theory without singularities.

6 Interacting Quantum Fields

This section provides a simple example of interacting quantum fields. If $X(n)$, $n = 0, 1, 2, \dots$, are operators on a Hilbert space, we define the *difference operator* ∇X by $\nabla X(n) = X(n+1) - X(n)$. We begin with a simple, but useful, lemma.

Lemma 6.1. *If $X(n)$ and $A(n)$ are operators on the same Hilbert space, $n = 0, 1, 2, \dots$, satisfying $\nabla X(n) = A(n)X(n)$, then*

$$\begin{aligned} X(n) &= [I + A(n-1)][I + A(n-2)] \cdots [I + A(0)] X(0) \\ &= \left[I + \sum_{j=0}^{n-1} A(j) + \sum \{A(j_1)A(j_2) : j_1, j_2 = 0, 1, \dots, n-1, j_2 < j_1\} \right. \\ &\quad + \sum \{A(j_1)A(j_2)A(j_3) : j_1, j_2, j_3 = 0, 1, \dots, n-1, j_3 < j_2 < j_1\} \\ &\quad \left. + \cdots + A(n-1)A(n-2) \cdots A(1)A(0) \right] X(0) \end{aligned}$$

Proof. We have that $X(n+1) - X(n) = A(n)X(n)$ so $X(n+1) = [I + A(n)]X(n)$. Replacing $n+1$ by n we obtain

$$\begin{aligned} X(n) &= [I + A(n-1)]X(n-1) = [I + A(n-1)][I + A(n-2)]X(n-2) \\ &= \cdots = [I + A(n-1)][I + A(n-2)] \cdots [I + A(0)]X(0) \end{aligned}$$

The second equality follows by induction. □

For $x \in \mathcal{S}_4$ define the field operators $\pi(x) = \phi(x) + \phi(x)^*$ on the Hilbert space \mathcal{H}_m as discussed in Section 5. Also, let $\sigma(x)$, $x \in \mathcal{S}_4$, be similarly defined field operators on the Hilbert space \mathcal{H}_M , $M \geq 0$. We think of $\pi(x)$ as describing π -particles of mass m and $\sigma(x)$ as describing σ -particles of mass M . The interaction between these two types of particles will be described on the tensor product $\mathcal{H} = \mathcal{H}_m \otimes \mathcal{H}_M$. Central roles are played by two operators on \mathcal{H} . These operators are functions of time $x_0 = 0, 1, 2, \dots$, and are the self-adjoint *interaction Hamiltonian* $H(x_0)$ and the *scattering operator* $S(x_0)$. We assume that $S(0) = I$ and that $S(x_0)$ satisfies the discrete Schrödinger's equation:

$$-i\nabla_{x_0} S(x_0) = H(x_0)S(x_0)$$

It follows from Lemma 6.1 that

$$\begin{aligned}
S(n) = I + i \sum_{j=0}^{n-1} H(j) + i^2 \sum \{H(j_1)H(j_2): j_1, j_2 = 0, 1, 2, \dots, n-1, j_2 < j_1\} \\
+ i^3 \sum \{H(j_1)H(j_2)H(j_3): j_1, j_2, j_3 = 0, 1, 2, \dots, n-1, j_3 < j_2 < j_1\} \\
+ \dots + i^n H(n-1)H(n-2) \dots H(0) \tag{6.1}
\end{aligned}$$

It is interesting that (6.1) resembles the inclusion-exclusion principle which is useful in probability and combinatorics theory. In the usual continuum theory, (6.1) has the form of a very complicated integral involving time-ordered products. In (6.1) the time-ordering is simpler and results in fewer terms. Essentially the only experiments available in quantum field theory are scattering experiments and the main objective in quantum field theory is to calculate the *final scattering operator* $S = \lim_{n \rightarrow \infty} S(n)$. Unfortunately, (6.1) cannot usually be solved in closed form to find S . The only thing we can do is use approximations or perturbative techniques. The interaction Hamiltonian is usually given in terms of the *Hamiltonian density* $K(x)$ by

$$H(x_0) = \sum \{K(x_0, x_1, x_2, x_3): \|x\|_4 \geq 0\}$$

In our particular example, suppose we consider the scattering of two π -particles interacting with a σ -particle. We take $K(x) = g\pi^2(x) \otimes \sigma(x)$ [8], where g is the *coupling constant*. When g is small, not as many terms are needed for approximations. Assume that the two π -particles initially have energy-momenta p and q giving rise, after scattering, to two π -particles with final energy-momenta p' and q' . Instead of writing the initial and final states as $|p\rangle \otimes |q\rangle$, $|p'\rangle \otimes |q'\rangle$ we use the simpler notation $|pq\rangle$ and $|p'q'\rangle$, respectively. The probability amplitude for the event of interest is $\langle p'q'|S|pq\rangle$ and the probability becomes

$$|\langle p'q'|S|pq\rangle|^2$$

Applying (6.1), the first two terms of $\langle p'q'|S|pq\rangle$ have the form

$$\langle p'q' | pq \rangle + i \sum_{\|x\|_4 \geq 0} \langle p'q'|K(x)|pq\rangle$$

Assuming that $|pq\rangle \neq |p'q'\rangle$, the two vectors are orthogonal so the first term is zero. The second term contains one $K(x)$ and hence only one $\sigma(x)$ field. This

applied to $|pq\rangle$ gives 0 for the annihilation part or a state of the form $|pq, r\rangle$ for the creation part. Since $\langle p'q' | pq, r\rangle = 0$ we again obtain 0. Similarly, any term with an odd number of $K(x)$ gives 0. The third term is nonzero and is treated in a similar way. However, there are quite a few possibilities and it appears that the best way to keep track is to employ Feynman diagrams [8]. We shall leave the details to later works.

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